

Snowmass2021 - Letter of Interest

## ***A Scintillating $n$ -type GaAs detector for sub-GeV Dark Matter Direct Detection***

### **Thematic Areas:**

- (CF1) Dark Matter: Particle Like
- (Other) [*IF1:Quantum Sensors*]

### **Contact Information:**

Rouven Essig (Stony Brook, rouven.essig@stonybrook.edu)

### **Authors:**

Karl Berggren (MIT), Edith Bourret-Courchesne (LBNL), Stephen Derenzo (LBNL), Rouven Essig (Stony Brook), Maurice Garcia-Sciveres (LBNL), Boris Korzh (JPL), Noah Kurinsky (FNAL), Sae Woo Nam (NIST, Colorado), Jamie Luskin (Maryland), Daniel McKinsey (Berkeley, LBNL) Matt Pyle (Berkeley), Matthew D Shaw (JPL), Emma Wollman (JPL), Tien-Tien Yu (Oregon)

**Abstract:** Sub-GeV dark matter can interact with electrons or nuclei in a  $n$ -type gallium arsenide (GaAs) target suitably doped with silicon and boron and produce 1.33 eV scintillation photons. These photons can be detected with a sensitive photodetector such as a Transition Edge Sensor (TES) or a Superconducting Nanowire Single-Photon Detector (SNSPD). This detection concept could be successfully realized with a high-mass ( $>10$  kg) target and thus probe smaller interactions between dark and ordinary matter than many other existing or proposed efforts to probe sub-GeV dark matter. In particular, no dark counts are expected, since the GaAs target has no afterglow and the 1.33 eV photons are naturally produced and easily measured without the need of an amplification mechanism (such as electric fields). We discuss the detection concept, the R&D required to realize this detection concept, and our plans.

# 1 Introduction and Physics Goals

A wide array of dark matter (DM) direct-detection ideas have been put forward to search for sub-GeV DM, see e.g. <sup>1-4</sup> and references therein. Here we highlight the role a scintillating  $n$ -type gallium arsenide (GaAs) target suitably doped with silicon and boron, coupled together with a sensitive photodetector such as a Transition Edge Sensor (TES) or a Superconducting Nanowire Single-Photon Detector (SNSPD), can play in this search. Such a detector could probe DM-electron scattering for DM masses above  $\sim 1$  MeV<sup>5;6</sup>, bosonic DM (dark photons, scalars, and axionlike particles) being absorbed by electrons for DM masses down to  $\mathcal{O}(\text{eV})$ <sup>7</sup>, DM-nucleus elastic scattering for masses  $\gtrsim 100$  MeV, and DM-nucleus scattering via the Migdal effect<sup>8;9</sup> for masses  $\gtrsim 1$  MeV. The potential sensitivity of such a search for DM-electron scattering is shown in Fig. 1.

$N$ -type GaAs(Si,B) produces scintillation photons when (1) an electronic excitation event leaves one or more holes in the valence band that are (2) trapped by acceptor atoms and (3) radiatively recombine with  $n$ -type donor band electrons. Preliminary studies show that the concentrations of  $n$ -type electrons, boron acceptors, and silicon complex acceptors are important in maximizing the luminosity and minimizing the trapping of valence band holes on non-radiative centers.

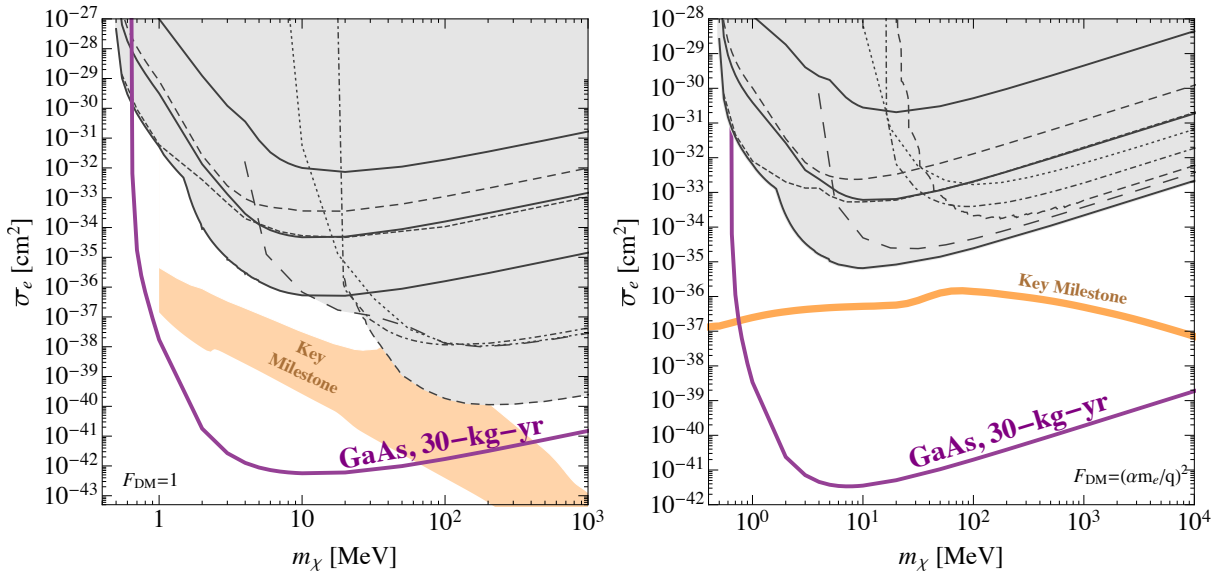


Figure 1: Projected sensitivity (thick magenta line) to dark-matter-electron scattering mediated by a heavy mediator (**left**) or light mediator (**right**) of a GaAs target assuming a 30 kg-year exposure. We assume zero background events for events with one or more photons, a radiative efficiency of 1, and a photon detection efficiency of 1. Existing constraints are shown in gray from SENSEI, DAMIC at SNOLAB, XENON10, XENON100, XENON1T, DarkSide-50, EDELWEISS, and CDMS-HV $\nu$ <sup>10-21</sup>. Orange regions labelled “Key Milestone” are from<sup>4</sup>.

## 2 Status and Required R&D

### 2.1 Gallium Arsenide Target Properties

$N$ -type GaAs suitably doped with silicon and boron is a luminous cryogenic scintillator<sup>22;23</sup> and has several properties that make it an excellent target for detecting sub-GeV DM:

- There are no naturally occurring radioactive isotopes of Ga or As.
- $N$ -type GaAs is commercially grown as 5 kg crystals.

- The donor electrons do not freeze out above a free-carrier concentration of  $8 \times 10^{15}/\text{cm}^3$ <sup>24</sup>.
- Metastable radiative centers that could cause afterglow are annihilated by delocalized donor electrons, as evidenced by the lack of thermally stimulated luminescence<sup>22</sup>.
- Luminosities  $>100$  photons/keV are observed without anti-reflective coatings.

Additional R&D and measurements are needed both to optimize the silicon and boron doping levels in the GaAs and to measure precisely the luminosity.

## 2.2 Photodetectors

The photodetectors must be able to measure 1.33 eV photons with negligible dark counts. Two natural possibilities are TES and SNSPD detectors. Their current status and the required R&D for realizing a large-mass GaAs DM detector are:

- **Transition Edge Sensors.** Very large area ( $45 \text{ cm}^2$ ) cryogenic photo detectors readout with TES based athermal phonon sensors have already achieved a measured 4 eV ( $\sigma_E$ ) sensitivity<sup>25</sup>. Since  $\sigma_E \propto \sqrt{A}$ , where A is the total instrumented area, this suggests that 600 meV resolutions in  $1 \text{ cm}^2$  detectors is already achievable. Additional improvements in the athermal phonon sensor design (lowering TES  $T_c$ , optimizing sensor geometry to improve athermal phonon collection efficiency) would allow one to further improve sensitivity and increase the instrumented area per channel past  $1 \text{ cm}^2$ .
- **Superconducting Nanowire Single-Photon Detector.** SNSPDs are the most advanced detectors available for time-resolved single photon counting from the UV to the infrared. They have demonstrated system detection efficiency as high as 98% at 1550 nm,<sup>26</sup> timing jitter below 3 ps,<sup>27</sup> energy thresholds as low as 0.125 eV, and dark count rates on the order of  $10^{-5}$  cps. SNSPDs have recently been demonstrated with mm-scale active areas, and typically have operating temperatures between 1 – 4 K. To support large-scale DM searches using GaAs targets, the active area of SNSPDs must be scaled to  $\text{cm}^2$  and beyond. This will require the development of new fabrication processes and techniques, and new on-chip multiplexing techniques, which allow the combination of signals from many nanowire sensor elements onto a single readout line.

The two proposed readout concepts (TES based athermal phonon sensors and SNSPDs) have vastly different susceptibilities to potential backgrounds and are thus complementary. Athermal phonon sensors are fundamentally sensitive to athermal phonons produced not only by photon absorption, but also by stress induced microfracture events and frictional rubbing with the mechanical support structure. To minimize these sources of environmental backgrounds, R&D is currently ongoing to suspend the devices from a double mass suspension system. As an additional protection against spurious dark counts, the GaAs crystal will also be instrumented with athermal phonon sensors; only events with coincident signals in both the GaAs and Ge photon detector would be indistinguishable from a DM interaction.

By contrast, SNSPDs have minimal sensitivity to spurious athermal phonons, since they are absorbed non-locally, while the absorption of a photon produces an extremely localized energy deposition that causes the SNSPD to transition. Thus, SNSPDs are not sensitive to potential phonon background sources. On the other hand, an SNSPD setup would not have a coincident athermal phonon signal on the GaAs target. As such, spurious single photon events, produced by very high energy photons travelling across an interface<sup>28</sup> or from higher energy nuclear recoils. It may be possible to mitigate such backgrounds by using multiphoton coincidence counting.

## 3 Plans

Several R&D steps are necessary to realize this detection concept:

- Optimize GaAs crystal size, surface roughness, and dopant density to maximize photon production and collection in the sensor.
- Athermal Phonon Sensor: Implement suspension system to suppress environmental vibrations.
- Athermal Phonon Sensor: Improve sensitivity by decreasing TES  $T_c$  and optimizing sensor geometry.
- SNSPD: increase active area from  $1 \text{ mm}^2$  to  $1 \text{ cm}^2$ .

## References

- [1] R. Essig et al., *Working Group Report: New Light Weakly Coupled Particles*, in *Proceedings, 2013 Community Summer Study on the Future of U.S. Particle Physics: Snowmass on the Mississippi (CsS<sup>2</sup>013): Minneapolis, Mn, Usa, July 29-August 6, 2013*, 2013, 1311.0029, <http://www.slac.stanford.edu/econf/C1307292/docs/IntensityFrontier/NewLight-17.pdf>.
- [2] J. Alexander et al., *Dark Sectors 2016 Workshop: Community Report*, 2016, 1608.08632, <http://lss.fnal.gov/archive/2016/conf/fermilab-conf-16-421.pdf>.
- [3] M. Battaglieri et al., *US Cosmic Visions: New Ideas in Dark Matter 2017: Community Report*, in *U.S. Cosmic Visions: New Ideas in Dark Matter College Park, MD, USA, March 23-25, 2017*, 2017, 1707.04591, <http://lss.fnal.gov/archive/2017/conf/fermilab-conf-17-282-ae-ppd-t.pdf>.
- [4] Report of the Workshop “Basic Research Needs for Dark Matter Small Projects New Initiatives” (October 2018) is available online at: [https://science.osti.gov/-/media/hep/pdf/Reports/Dark\\_Matter\\_New\\_Initiatives\\_rpt.pdf](https://science.osti.gov/-/media/hep/pdf/Reports/Dark_Matter_New_Initiatives_rpt.pdf).
- [5] S. Derenzo, R. Essig, A. Massari, A. Soto and T.-T. Yu, *Direct Detection of sub-GeV Dark Matter with Scintillating Targets*, *Phys. Rev.* **D96** (2017) 016026, [1607.01009].
- [6] R. Essig, J. Mardon and T. Volansky, *Direct Detection of Sub-GeV Dark Matter*, *Phys. Rev.* **D85** (2012) 076007, [1108.5383].
- [7] I. M. Bloch, R. Essig, K. Tobioka, T. Volansky and T.-T. Yu, *Searching for Dark Absorption with Direct Detection Experiments*, *JHEP* **06** (2017) 087, [1608.02123].
- [8] M. Ibe, W. Nakano, Y. Shoji and K. Suzuki, *Migdal Effect in Dark Matter Direct Detection Experiments*, *JHEP* **03** (2018) 194, [1707.07258].
- [9] R. Essig, J. Pradler, M. Sholapurkar and T.-T. Yu, *Relation between the migdal effect and dark matter-electron scattering in isolated atoms and semiconductors*, *Phys. Rev. Lett.* **124** (Jan, 2020) 021801.
- [10] R. Essig, A. Manalaysay, J. Mardon, P. Sorensen and T. Volansky, *First Direct Detection Limits on Sub-GeV Dark Matter from Xenon10*, *Phys. Rev. Lett.* **109** (2012) 021301, [1206.2644].
- [11] R. Essig, T. Volansky and T.-T. Yu, *New Constraints and Prospects for Sub-GeV Dark Matter Scattering Off Electrons in Xenon*, *Phys. Rev.* **D96** (2017) 043017, [1703.00910].
- [12] XENON10 collaboration, J. Angle et al., *A Search for Light Dark Matter in Xenon10 Data*, *Phys. Rev. Lett.* **107** (2011) 051301, [1104.3088].
- [13] XENON collaboration, E. Aprile et al., *Low-Mass Dark Matter Search Using Ionization Signals in Xenon100*, *Phys. Rev.* **D94** (2016) 092001, [1605.06262].
- [14] DARKSIDE collaboration, P. Agnes et al., *Constraints on Sub-GeV Dark-Matter Electron Scattering from the DarkSide-50 Experiment*, *Phys. Rev. Lett.* **121** (2018) 111303, [1802.06998].
- [15] SUPERCDMS collaboration, R. Agnese et al., *First Dark Matter Constraints from a Supercdms Single-Charge Sensitive Detector*, *Phys. Rev. Lett.* **121** (2018) 051301, [1804.10697].
- [16] SENSEI collaboration, M. Crisler, R. Essig, J. Estrada, G. Fernandez, J. Tiffenberg, M. Sofo haro et al., *Sensei: First Direct-Detection Constraints on Sub-GeV Dark Matter from a Surface Run*, *Phys. Rev. Lett.* **121** (2018) 061803, [1804.00088].

- [17] SENSEI collaboration, O. Abramoff et al., *Sensei: Direct-Detection Constraints on Sub-GeV Dark Matter from a Shallow Underground Run Using a Prototype Skipper-Ccd*, *Phys. Rev. Lett.* **122** (2019) 161801, [1901.10478].
- [18] DAMIC collaboration, A. Aguilar-Arevalo et al., *Constraints on Light Dark Matter Particles Interacting with Electrons from Damic at Snolab*, *Phys. Rev. Lett.* **123** (2019) 181802, [1907.12628].
- [19] SENSEI collaboration, L. Barak et al., *SENSEI: Direct-Detection Results on sub-GeV Dark Matter from a New Skipper-CCD*, 2004.11378.
- [20] EDELWEISS collaboration, Q. Arnaud et al., *First germanium-based constraints on sub-MeV Dark Matter with the EDELWEISS experiment*, 2003.01046.
- [21] SUPERCDMS collaboration, D. Amaral et al., *Constraints on low-mass, relic dark matter candidates from a surface-operated SuperCDMS single-charge sensitive detector*, 2005.14067.
- [22] S. Derenzo, E. Bourret, S. Hanrahan and G. Bizarri, *Cryogenic Scintillation Properties of n-Type GaAs for the Direct Detection of  $\text{MeV}/c^2$  Dark Matter*, *J. Appl. Phys.* **123** (2018) 114501, [1802.09171].
- [23] S. Vasiukov, F. Chiossi, C. Braggio, G. Carugno, F. Moretti, E. Bourret et al., *GaAs as a Bright Cryogenic Scintillator for the Detection of Low-Energy Electron Recoils From  $\text{MeV}/c^2$  Dark Matter*, *IEEE Trans. Nucl. Sci.* **66** (2019) 2333–2337.
- [24] M. Benzaquen, D. Walsh and K. Mazuruk, *Conductivity of n-type gaas near the mott transition*, *Phys. Rev. B* **36** (Sep, 1987) 4748–4753.
- [25] SUPERCDMS collaboration, I. Alkhatib et al., *Light Dark Matter Search with a High-Resolution Athermal Phonon Detector Operated Above Ground*, 2007.14289.
- [26] D. Reddy, A. Lita, S. W. Nam, R. Mirin and V. Verma, *Achieving 98% system efficiency at 1550 nm in superconducting nanowire single photon detectors*, in *Proceedings, 2019 Rochester Conference on Coherence and Quantum Optics (CQO-11), Rochester, NY, USA, August 4 - August 8, 2019*, 2019, <https://www.osapublishing.org/abstract.cfm?uri=CQO-2019-W2B.2>.
- [27] B. Korzh, Q.-Y. Zhao et al., *Demonstration of sub-3 ps temporal resolution with a superconducting nanowire single-photon detector*, *Nature Photonics* **14** (2020) 250–255, [1804.06839].
- [28] T. F. Silva, A. L. Bonini, R. R. Lima, N. L. Maidana, A. A. Malafronte, P. R. Pascholati et al., *Optical transition radiation used in the diagnostic of low energy and low current electron beams in particle accelerators*, *Review of Scientific Instruments* **83** (2012) 093301.