

1 Snowmass2021 - Letter of Interest

2 *HAYSTAC – Pioneering the Quantum Frontier*

3 **Thematic Areas:** (check all that apply /■)

4 ■ (IF1) Quantum Sensors

5 ■ (CF2) Dark Matter: Wavelike

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9 Collaboration (optional): HAYSTAC Collaboration

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12 **Abstract:**

13 HAYSTAC is a microwave cavity dark matter axion search designed as an innovation testbed and data
14 pathfinder for the $15 - 50 \mu\text{eV}$ ($3.5 - 12 \text{ GHz}$) mass range. Commissioned in 2014 with a single Josephson
15 Parametric Amplifier (JPA) receiver, it has operated from the beginning with a system noise temperature
16 N_{sys} within an order of magnitude of the Standard Quantum Limit N_{SQL} . Recently the experiment has
17 run with a dual JPA squeezed state receiver circumventing the standard quantum limit entirely, achieving
18 a sensitivity to axions at $1.38 \times \text{KSVZ}$ despite its small volume $V = 1.5 \text{ L}$. Breaking the quantum barrier
19 is a significant milestone; not only is this the first quantum-enhanced dark matter experiment ever, but to-
20 gether with LIGO it is only one of two experiments to have employed squeezed states to hunt for new
21 fundamental physics. Parallel developments in microwave cavities for this frequency range have included
22 tunable Photonic Band Gap structures, and more recently metamaterial plasmonic resonators. While main-
23 taining its identity as a technology test bed, future operation will emphasize production running in this mass
24 range, where a much more ambitious receiver is under development (the CEASEFIRE project) based on a
25 squeezing and state-swapping protocol in order to best probe the wide area of open axion parameter space.

26 **Science landscape:** Since the last Snowmass in 2013, dramatic progress has been made in the fields of
27 both axion theory and experiment, and the axion has emerged as one of the two leading Dark Matter (DM)
28 candidates. From the perspective of axion cosmology, we need to be prepared to search over a much wider
29 range of masses than is currently possible experimentally as some inflationary scenarios naturally allow the
30 axion mass to be much lower than previously thought, as low as 10^{-12} eV^{1;2}. On the other hand, for post-
31 inflation PQ-symmetry breaking, predictions of the axion mass corresponding to a dark matter fraction of
32 $\Omega_{DM} \approx 0.27$ have tightened up dramatically, with the $m_a \approx 15\text{-}35$ μeV range now being a particularly well
33 motivated region^{3;4}. Similarly, it has been recently realized that ALPogenesis can simultaneously explain
34 both baryon and dark matter densities, but with photon couplings orders of magnitude stronger than for the
35 standard QCD axion⁵. Taken together, this suggests a broad search philosophy, namely that one should
36 cover as much of the currently available parameter space as rapidly as possible, with an eye towards an
37 early discovery. On the experimental side, in the past several years, three experiments have now published
38 results probing the QCD model band in the $> \mu\text{eV}$ range: Haloscope At Yale Sensitive To Axion Cold
39 dark matter (HAYSTAC)⁶⁻⁸, ADMX⁹, and CAPP¹⁰. Additionally, major experiments are currently under
40 development to search for axions of lower (ABRACADABRA¹¹, DM Radio¹²⁻¹⁴, CASPER¹⁵) and higher
41 masses (MADMAX¹⁶).

42 **Breaking the quantum barrier:** Over the past decade, the HAYSTAC experiment has been steadily
43 pushing the limits of quantum measurement in the search for DM axions. First funded by NSF in 2011
44 and later by the Heising-Simons Foundation, the HAYSTAC experiment — a collaboration of Yale, UC
45 Berkeley, Colorado, and now Johns Hopkins — was designed to be both an innovation testbed and a data
46 pathfinder in the 4–12 GHz range. The experiment was intended to be an agile technology platform where
47 the latest advances in quantum measurement could be explored along with novel concepts for microwave
48 cavities which satisfy the stringent requirements of an axion search while maintaining high figures of merit.
49 From its inaugural data run in 2015, HAYSTAC operated with a system noise level N_{sys} within a factor of
50 2-3 of the standard quantum limit N_{SQL} , and is to date still the only microwave cavity axion experiment to
51 come within an order of magnitude of N_{SQL} ^{6;7;17}. Going further, however; we have just recently submitted
52 for publication our first run with a Squeezed State Receiver (SSR) which manipulates quantum states in
53 order to circumvent the standard quantum limit entirely⁸. Breaking the quantum barrier is a significant
54 milestone; not only is this the first quantum-enhanced DM experiment ever, but together with LIGO it is
55 only one of two experiments to have ever employed squeezed states to search for new fundamental physics.
56 Squeezing with other optimizations¹⁸ have accelerated our search by a factor of ≈ 3 over the standard mode
57 of operation. Coupled with additional experimental improvements, for our next data runs we seek to improve
58 the scan rate by an additional factor of 4. In addition, thanks to a radical new design being developed by our
59 Colorado/JILA collaborators we have the potential to increase this scan rate even further.

60 **Innovative data analysis strategies:** In addition, HAYSTAC has led a push in the axion direct detection
61 community for improved data processing and statistical analysis. In 2017¹⁹, HAYSTAC standardized the
62 use of maximum-likelihood estimation throughout the data processing and optimized the spectral filtering
63 used across all haloscope searches. In doing so, we provided the first codification of haloscope analysis
64 protocols since 2001²⁰, during which time these protocols have evolved significantly. Practices pioneered by
65 HAYSTAC have since been adopted by all other QCD axion-sensitive haloscopes^{10;21}. In 2020, HAYSTAC
66 optimized the framework commonly used for statistical inference in axion direct detection, allowing for
67 conservative results reported with 30-50% faster effective scan rates, and more flexible stopping criteria,
68 which aided our most recent data run^{8;18}. This novel framework has since been used advantageously in an
69 ALP DM search conducted using existing data from the JILA eEDM experiment and extending 17 orders of
70 magnitude lower in mass than the QCD axions that HAYSTAC targets²².

71 **Future plans:** While we will continue to advance technology both on the receiver and cavity fronts, at

72 the same time HAYSTAC will now additionally function as a data production experiment aiming to search
73 much of the 15-30 μeV (3.5-8.0 GHz) range. Ideally, the experiment will upgrade to a Nb_3Sn magnet as
74 soon as possible, improving the scan rate $R \propto B^2V$ of 10 and enable coverage of the 30-50 μeV (8.0-12
75 GHz) range. While there are multiple paths to quantum enhancement²³, we argue that squeezing is by far
76 the most practical and feasible in the near term - indeed it has already delivered science results. The current
77 SSR design is broadly tunable and has a clear path for continuous improvements. While the collaboration
78 now comprises four institutions with Johns Hopkins (where former HAYSTAC postdoc D. H. Speller is now
79 faculty), expanding by one new group may be warranted to ensure adequate depth in both operations and
80 analysis, for a data production and technology R&D experiment.

81 **State of the R&D:** The question should be asked whether extensions of the present SSR, and a possible
82 next generation squeezer of greatly enhanced capability and its incorporation into the HAYSTAC experiment
83 is well supported by both the technology and the R&D. Concerning the current platform, the Josephson Para-
84 metric Amplifiers (JPAs) and microwave cavities for the 3.5-8.0 GHz phase are already in hand and ready
85 for deployment. For the 8.0-12.0 GHz operation, the current JPA design is satisfactory; but some of the con-
86 ventional microwave components in the receiver will need to be modified for use at higher frequencies. In
87 regard to the microwave cavities, extensions of the current design²⁴ look promising and R&D is underway.
88 The Berkeley group has studied tunable resonators based on Photonic Band Gap structures to eliminate the
89 forest of intruder TE modes which mix with the TM modes of interest, eroding frequency coverage. This
90 has been the bane of all cavity schemes to extend their dynamical range upward in frequency. The state of
91 cavity R&D has already produced several interesting results with engineering prototypes of actual cavities
92 to be manufactured and tested in the coming year.

93 The one R&D program which is considered moderate risk is the development of a next-generation squeez-
94 ing and state swapping protocol to achieve very large squeezing factors currently limited by losses in the
95 circulator. The Colorado group is therefore pursuing an architecture that does not require a circulator, under
96 the banner of the CEASEFIRE project²⁵ with a bench demonstration that could be ready in two-four years
97 and integration into HAYSTAC soon thereafter.

98 **Yale's Initiative in Quantum Science:** Quantum science and advanced instrumentation development
99 for fundamental research have been identified as the top priorities in Yale's science initiative¹. Yale Uni-
100 versity is also one of the leading partners in the recently announced DOE "Co-design Center for Quantum
101 Advantage"². Over the coming years, Yale will be developing a Physics Science and Engineering Building
102 adjacent to the current Wright Laboratory including a state-of-the-art Advanced Instrumentation Develop-
103 ment Center, fabrication facilities, cleanrooms and technical staff to support Yale's initiative in this field³.
104 The building will provide new laboratory space for the HAYSTAC experiment by 2023 and enable research
105 and production on quantum devices. Together with the historic strength of the Yale quantum sciences group
106 this will represent a focal point for the East Coast comparable to NIST Boulder. The new HAYSTAC lab will
107 have a closed-loop LHe system supporting all low-temperature operations in the lab, including the Nb_3Sn
108 magnet we hope to procure for HAYSTAC. This development represents an opportunity for Yale to become
109 a national center for quantum-enhanced dark matter research. This new facility also represents an excellent
110 future resource for all upcoming HAYSTAC runs.

111 **Long-range R&D:** HAYSTAC will continue to promote and follow the R&D on a quantum non-
112 demolition Rydberg atom single-quantum detector, being led by R. Maruyama (as discussed in the LOI
113 submitted to IF1/CF2). While this is potentially an exceedingly powerful strategy for axion searches, this is
114 long-range R&D, and the pathway for integration into HAYSTAC remains an open question.

¹University Science Strategy Committee Report

²"Yale scientists to help lead national quantum center" - Yale News August 2020

³University Wide Academic Priorities

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