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The Microwave SQUID Multiplexer for Cosmology and Cryogenic Particle Detection

Thematic Areas:

■ (IF1) Quantum Sensors

- (IF2) Photon Detectors
- (IF6) Calorimetry

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Abstract: Cryogenic sensors have found a large range of applications for astroparticle detection. Due to integration complexity and thermal loading from cryogenic wiring, the ability to read out multiple detectors on a single wire with cryogenic multiplexing technologies with minimal readout noise penalty is of utmost importance as experiments are scaled to ever larger detector counts. The microwave SQUID multiplexer (μ mux) couples an incoming detector signal to a unique GHz-frequency resonance, thus combining the multiplexability of MKIDs with the clean separation of detection and readout interfaces. This enables multiplexing factors up to two orders of magnitude larger than conventional cryogenic multiplexing schemes. The wide frequency operation span enables large detector counts for low-bandwidth bolometric applications such as CMB cosmology while maintaining clean interfaces between the detection and readout schemes. Additionally, the large frequency bandwidth and fast resonator response allow for cryogenic particle detection, such as low-mass threshold dark matter searches, beta decay end point measurements to determine the lightest neutrino mass, and coherent elastic neutrino-nucleon scattering. Recent developments in warm electronics have enabled rapid progress in μ mux development and maturity across several application areas.

Introduction Superconducting and/or cryogenic sensors have become widespread in astroparticle detection thanks to developments in repeatable fabrication, availability of commercial refrigeration tools, and technological superiority for several applications. Particularly popular examples are transition-edge sensors¹² and magnetic micro-calorimeters³, which have been deployed in experiments measuring WIMP-like dark matter⁴, millimeter and sub-mm wavelength photons⁵, x-ray and gamma-ray spectra⁶⁷, and neutrinos⁸, among others.

The multiplexer works by means of inductively coupling the incoming sensor signal to an rf SQUID, which is in turn inductively coupled to a unique microwave frequency resonator. Changes in flux threading the SQUID loop shift the frequency of the microwave resonance. A common flux ramp line is coupled to all of the SQUIDs and ramped much faster than the individual sensor signal, thus linearizing the SQUID response and eliminating the need for individual feedback lines. This flux ramp has the additional benefit of transforming the incoming detector signal into a phase modulation of the SQUID response. Thus, a single microwave transmission line on a pair of coaxial cables may read out thousands of individual sensors across a typical 4-8GHz bandwidth⁹¹⁰.



Figure 1: A circuit diagram for μ mux used with transition-edge sensors (TESs). The working principle is similar for magnetic microcalorimeters (MMCs), except the dc-biased TESs are replaced with individual MMCs.

For mm and sub-mm photon detection applications, typical Q factors achieved and generous spacing between resonators to avoid frequency collisions allow for O(4000) channels to be multiplexed on a single readout line. This represents almost two orders of magnitude improvement over typical time-domain or MHz-frequency domain multiplexing schemes presently in use.¹¹¹² For cryogenic particle detection applications, the large operating bandwidth allows for large electrical bandwidth and fast response.

Recent Development Recent developments in warm rf electronics have enabled rapid progress in the maturity of μ mux. In particular, the advent of tone-tracking electronics which "follow" the resonance dip as it is modulated, ensuring that the input microwave excitation tone is always attenuated by the 10-20dB resonance dip. This development, combined with advances in the linearity performance of cryogenic amplifiers, has opened a pathway for μ mux to reach its full multiplexing potential¹³.

In the austral summer of 2018/19, a single receiver of the Keck Array, part of the BICEP/Keck series of CMB polarimeters, was retrofitted with a μ mux 150GHz focal plane and deployed to the South Pole for the 2019 observing season¹⁵. The μ mux electronics were integrated with existing time-domain multiplexing (TDM) electronics operating on other receivers, timing, and telescope control/DAQ systems to produce an end-to-end demonstration of μ mux for CMB. The on-sky performance in optimized readout channels

was found to be comparable to the performance of the same TES detectors operating with time-domain multiplexing in a previous deployment, allowing for a direct comparison of readout technologies.

The compatibility of the dc-biasing for TESs between TDM and μ mux allow for minimal changes to the focal plane package architecture, while the microwave plumbing allows for overall simplification and massive reduction of the readout wiring, making μ mux an attractive upgrade for existing TDM systems seeking to expand their multiplexing capability without a total experimental redesign. This additionally allows for the separate optimization, fabrication, and characterization of the detector and readout elements, allowing for large scale production for big experiments.

Future CMB experiments that intend to use μ mux include the BICEP Array high frequency receiver, the Ali-CPT telescope, and the Simons Observatory, each of which intend to deploy over 20,000 TESs at mK temperatures in a single cryostat. A significant amount of development has gone into the large-scale packaging, wiring, and integration of μ mux for each of these experiments.¹⁶¹⁷ Further development is underway to explore repeatable packaging of μ mux devices while maintaining good rf transmission.

The large bandwidth and fast response of μ mux makes it an attractive option for calorimetry applications in addition to the aforementioned developments in bolometry. μ mux systems are being developed for a number of terrestrial and space-based x-ray calorimetry experiments, including beamline spectrometers and the Lynx x-ray satellite mission^{7 18 19}. In addition to TES-based calorimeters, μ mux has additionally been developed for magnetic microcalorimeters²⁰, particularly for direct neutrino mass measurements with the end point energy of ¹⁶³Ho as in HOLMES or ECHo^{21 22}. The need for multiplexing arises not only from the need for larger statistics but also the desire to avoid pileup at any given detector. Here, the key advantage of μ mux lies in the ability to better resolve the pulse shape of an incoming photon since it is not limited by the switching time of the SQUID as in time- or code- division multiplexing systems or the complications of putting the inductor in the bias circuit as in MHz-frequency domain multiplexing systems.

Future Applications The recent rapid progress of μ mux demonstrates its promise for cosmology and particle physics applications. The large readout bandwidth and fast response make it particularly attractive for cryogenic particle detection experiments, while the high multiplexing factor and minimal noise penalty allow for a wide variety of cosmological applications. Renewed interest in microwave resonator-based experimental techniques thanks to recent developments in quantum computing architectures additionally have contributed to the availability of rf parts and expertise.

Upcoming CMB experiments will set tight constraints on primordial gravitational waves, light relics, the sum of the neutrino masses and the neutrino mass hierarchy, the universe's expansion history and dark energy, and several dark matter scenarios, among other cosmological parameters of interest. As these experiments seek to deploy tens or even hundreds of thousands of cryogenic detectors in order to maximize sensitivity, the multiplexing advantage of μ mux becomes ever more apparent. Spectroscopic measurements of cosmic dawn and other mm- and sub-mm wave experiments similarly benefit from low-noise cryogenic multiplexing at large scales. To truly harness the multiplexing benefit and make large arrays possible at scalable cost and timeline, the parallel development of fabrication and characterization of large numbers of TESs and microwave SQUIDs, as well as installation and calibration resources will be necessary.

TES and MMC-based cryogenic sensors for ultra-sensitive particle detection are enabled by μ mux, which offers large readout bandwidth and relatively high multiplexing factors with fast response. These include end point measurements for direction measurement of the lightest neutrino mass, coherent elastic neutrino-nucleon scattering measurements, and light dark matter direct detection. The cryogenic particle detection community stands to benefit massively from the relative maturity of photon detector-based focal plane multiplexing technologies, presenting a viable path to distributed massive non-focal plane experiments with thousands or even tens of thousands of detectors.

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