

1 Snowmass2021 - Letter of Interest

2 *Prospects for keV Sterile Neutrino Searches with* 3 *KATRIN*

4 **NF Topical Groups:** (check all that apply /■)

5 (NF1) Neutrino oscillations

6 ■ (NF2) Sterile neutrinos

7 ■ (NF3) Beyond the Standard Model

8 (NF4) Neutrinos from natural sources

9 ■ (NF5) Neutrino properties

10 (NF6) Neutrino cross sections

11 (NF7) Applications

12 (TF11) Theory of neutrino physics

13 (NF9) Artificial neutrino sources

14 (NF10) Neutrino detectors

15 (Other) [*Please specify frontier/topical group(s)*]

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19 Collaboration: KATRIN

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21 **Authors:** The full author list follows the text and references.

22 **Abstract:** Right-handed neutrinos are a minimal extension of the Standard Model of Particle Physics which
23 leads to the existence of new neutrino mass eigenstates with arbitrary mass scale. Neutrinos with masses in
24 the keV range and very weak mixing with active species form viable dark matter candidates. A promising
25 model-independent way to search for keV-scale sterile neutrinos is via high-precision beta spectroscopy,
26 where these particles manifest themselves as tiny distortion of the spectral shape. The Karlsruhe Tritium
27 Neutrino (KATRIN) experiment, equipped with a novel multi-pixel silicon-drift-detector focal plane array
28 and read-out system, named the TRISTAN detector, has the potential to surpass the sensitivity of previous
29 laboratory-based searches.

30 As prominent dark matter candidates, sterile neutrinos in the keV mass range are at the crossroads of
31 particle physics, astrophysics and cosmology [1, 2]. Depending on their production mechanism in the early
32 universe, they can act as effectively cold, cool, or even warm dark matter, which in turn have drastically
33 different influences on the structure formation of our cosmos [3]. A sizable fraction of warm dark matter
34 can mitigate tensions between observations and predictions of purely cold-dark-matter scenarios [4].

35 Both cosmological and astrophysical data yield strong bounds on keV-scale sterile neutrinos, but these
36 rely on an underlying cosmological model [5]. Applying such bounds assumes that all dark matter is com-
37 posed of sterile neutrinos, that a certain production mechanism was realized in the early universe, and
38 that the distribution of dark matter and the astrophysical foregrounds are understood. Therefore, model-
39 independent, laboratory-based experiments are indispensable to directly probe the existence and properties
40 of these hypothetical particles.

41 The current limits from laboratory-based searches are several orders of magnitude weaker than astro-
42 physical and cosmological limits. Yet, through renewed interest in the field, many promising proposals to
43 push the sensitivity of laboratory-based experiments have been put forward in recent years [6, 7, 8, 9]. One
44 of the most feasible approaches, to be realized in the near future, is to search for sterile neutrinos in beta de-
45 cays. The sterile neutrino, emitted in a beta-decay, would lead to a small, but characteristic distortion of the
46 beta-decay spectrum at an energy $E = E_0 - m_s$, where E_0 is the kinematic endpoint and m_s corresponds
47 to the sterile-neutrino mass. The main challenge of a keV-sterile neutrino search in beta decays is to reduce
48 both the statistical and systematic uncertainties to the sub-ppm level, in order to probe parameter space of
49 interest for particle physics and cosmology.

50 A world-leading beta-spectroscopy experiment is the KATRIN experiment [10]. The main objective of
51 KATRIN is the determination of the absolute neutrino-mass scale [11] with an unprecedented sensitivity of
52 200 meV, by measuring the tritium beta decay spectrum close to its endpoint in an integral way [10]. Yet,
53 its ultra-high source activity and excellent spectral analysis quality allow it to extend its physics program to
54 search for a wide variety of beyond-standard-model phenomena [12], including sterile neutrinos in a wide
55 mass range.

56 The targeted sensitivity for keV-scale sterile neutrino admixture is 10^{-6} . From a statistical point of view
57 the enormous source activity of KATRIN allows to reach mixing angles at the 10^{-8} level with approx. 1 year
58 of measurement time. However, reducing systematic uncertainties to the ppm-level poses a major challenge.

59 A major technological challenge arises when operating KATRIN to search for keV-scale sterile neutri-
60 nos. The signal of the new particle would appear far away from the endpoint, where the rate of tritium-decay
61 electrons is high. Accordingly, the counting rate at the detector will be many orders of magnitude higher
62 than in normal KATRIN operation, where only the spectral endpoint region is observed.

63 KATRIN is presently pursuing a research and development program into an advanced detector and read-
64 out system known as TRISTAN [13]. TRISTAN is built of Silicon Drift Detectors (SDDs), an ideally suited
65 technology designed for high-rate and high-resolution applications [14]. The upgraded detector is expected
66 to handle high rates (up to 10^9 cps) while at the same time providing excellent energy resolution (300 eV
67 FWHM @ 18 keV for electrons) and energy linearity (ppm-level). High-precision beta spectroscopy is
68 a novel application for SDDs, which are typically used for x-ray spectroscopy. To improve response to
69 electrons, we are exploring new production techniques to reduce the SDD entrance window thickness to less
70 than 30 nm.

71 New technology will be needed to integrate the 20-cm-diameter TRISTAN focal-plane array, with more
72 than 1000 channels, into ultra-high vacuum and high magnetic- and electric-field conditions. An advanced
73 data acquisition system, still located in the high-field regions, will perform a full waveform digitization. The
74 data, at rates exceeding Gb/s, will be sent via optical links to a high-performance FPGA system (situated in

75 the low-field regions) to perform event-by-event energy filtering and multiplicity analysis.

76 The TRISTAN detector system will complement the standard integral measurement of KATRIN (in
77 which the detector only counts electrons) with a high-resolution differential measurement (in which the
78 detector determines the energy of each electron). This novel idea of combining differential and integral
79 measurement modes will be a key to ruling out large classes of systematic uncertainties.

80 Over the past years major advances with respect to the keV-scale sterile neutrino search with KA-
81 TRIN have been made. An important milestone for the keV-scale sterile neutrino program of KATRIN
82 was achieved with the first tritium measurement in 2018 [15]. Thanks to a rather low tritium activity during
83 this commissioning campaign, an extended measurement interval was possible, which allowed the search
84 for sterile neutrinos in a mass range of up to about 1.6 keV. An excellent agreement of the undistorted
85 spectrum calculation to the data could be demonstrated in this broad energy range, improving the current
86 laboratory-based limits in the sterile neutrino mass range between 0.1 and 0.5 keV to the level of 10^{-3} [16].

87 To extend the mass range and improve the sensitivity further, the novel TRISTAN detector system will be
88 necessary. With the first generation of prototype detectors, the excellent performance of SDDs [13] and their
89 applicability for beta-spectroscopy [17] was demonstrated. The development of the first TRISTAN detector
90 modules has been completed in the beginning of 2020. A TRISTAN module comprises 166 detector pixels
91 with an integrated read-out amplification. A fully equipped detector module is planned to be installed at the
92 KATRIN monitor spectrometer [18] in 2021. The final detector system will be integrated into the KATRIN
93 beamline in a staged approach. Phase 1 is defined as a 9-module system, which requires minimal changes to
94 the current KATRIN focal-plane detector chamber. Phase 2 is a further upgrade to 21 modules, for which a
95 new detector chamber with an optimized electromagnetic field configuration is required. All upgrades will
96 take place after the successful completion of the neutrino-mass measurement program of KATRIN.

97 **Acknowledgments** We acknowledge the support of Helmholtz Association, Ministry for Education and
98 Research BMBF (5A17PDA, 05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3), Helmholtz Alliance
99 for Astroparticle Physics (HAP), Helmholtz Young Investigator Group (VH-NG-1055), Max Planck Re-
100 search Group (MaxPlanck@TUM), and Deutsche Forschungsgemeinschaft DFG (Research Training Groups
101 GRK 1694 and GRK 2149, Graduate School GSC 1085 - KSETA, and SFB-1258) in Germany; Min-
102 istry of Education, Youth and Sport (CANAM-LM2011019, LTT19005) in the Czech Republic; and the
103 United States Department of Energy through grants DE-FG02-97ER41020, DE-FG02-94ER40818, DE-
104 SC0004036, DE-FG02-97ER41033, DE-FG02-97ER41041, DE-AC02-05CH11231, DE-SC0011091, and
105 DE-SC0019304, and the National Energy Research Scientific Computing Center. This project has also
106 received funding from the European Research Council (ERC) under the European Union Horizon 2020
107 research and innovation program (grant agreement No. 852845).

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