Snowmass2021 - Letter of Interest

The JUNO-TAO Experiment

NF Topical Groups: (check all that apply \Box/\blacksquare)

□ (NF1) Neutrino oscillations

- (NF2) Sterile neutrinos
- (NF3) Beyond the Standard Model
- \Box (NF4) Neutrinos from natural sources
- \Box (NF5) Neutrino properties
- \Box (NF6) Neutrino cross sections
- (NF7) Applications
- \Box (NF8) Theory of neutrino physics
- (NF9) Artificial neutrino sources
- (NF10) Neutrino detectors
- \Box (Other) [*Please specify frontier/topical group(s*)]

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Abstract: The Taishan Antineutrino Observatory (TAO, also known as JUNO-TAO) is a satellite experiment of the Jiangmen Underground Neutrino Observatory (JUNO). The experiment consists of a ton-level liquid scintillator detector placed at ~ 30 m from a 4.6 GW_{th} reactor core of the Taishan Nuclear Power Plant. The main goal is to measure the reactor antineutrino spectrum with sub-percent energy resolution, providing a reference spectrum for JUNO as well as a benchmark for nuclear databases and other experiments. The detector design consists of a spherical acrylic vessel containing 2.8 ton gadolinium-doped liquid scintillator viewed by 10 m² Silicon Photomultipliers (SiPMs) with $\sim 50\%$ photon detection efficiency and providing around 95% photon coverage. The photoelectron yield will be around 4500 per MeV, an order of magnitude higher than any existing large-scale liquid scintillator detector. The detector will operate at -50°C to mitigate the impact of SiPM dark noise and will be well shielded from cosmogenic backgrounds and ambient radioactivity. About 2000 reactor antineutrinos will be collected per day with an expected background-to-signal ratio of $\sim 10\%$. Operations are expected to begin as soon as 2022.

Key words: reactor antineutrino spectrum, JUNO-TAO

Physics Motivation

The three-neutrino oscillation paradigm is well supported by experiments using solar neutrinos, atmospheric neutrinos, accelerator neutrinos and reactor antineutrinos. The three neutrino mixing angles, θ_{12} , θ_{13} , θ_{23} , as well as the two independent mass-squared splittings, $\Delta m_{21}^2 \equiv m_2^2 - m_1^2$, $|\Delta m_{31}^2| \equiv |m_3^2 - m_1^2|$ (or $|\Delta m_{32}^2|$), where m_1, m_2 , and m_3 are the three mass eigenvalues, have been measured to few-percent precision. However, even assuming neutrinos are Dirac fermions, the Dirac CP-violation phase and the neutrino mass ordering (NMO) are still unknown. The Jiangmen Underground Neutrino Observatory (JUNO) experiment in South China aims to resolve the NMO and further improve the measurements of θ_{12} , Δm_{21}^2 , and $|\Delta m_{31}^2|$ to sub-percent precision by measuring the oscillated energy spectrum of the antineutrinos emitted by eight reactors from the Yangjiang and Taishan nuclear power plants at a distance of 53 km. To meet these two major goals, the JUNO detector is designed to reach an energy resolution of $3\%/\sqrt{E[MeV]}$, where *E* is the visible energy [1]. Recent experiments such as Daya Bay have shown that the predicted antineutrino spectrum disagrees with observations [2]. Moreover, it has been speculated that unknown fine structure in the energy spectrum [3, 4] could affect the sensitivity of JUNO experiment [5, 6]. In order to mitigate this possibility, a satellite experiment called the Taishan Antineutrino Observatory (JUNO-TAO) has been proposed to provide a reference spectrum to JUNO [7].

The Taishan Nuclear Power Plant is located in the Chixi town of Taishan city in the Guangdong province. This plant currently hosts two European Pressurised Reactors (EPR) in operation, each with 4.6 GW of thermal power. JUNO-TAO is located in a basement 9.6 m underground, outside of the concrete containment shell of one of the two reactors. The primary goal of JUNO-TAO is to eliminate any possible model dependence in JUNO's oscillation measurements by providing a reference spectrum. Although a $3\%/\sqrt{E[MeV]}$ energy resolution would be sufficient to accomplish this goal, it is advantageous to achieve a higher energy resolution to study the fine structure of the reactor antineutrino spectrum and produce a state-of-the-art measurement that could uncover unexpected features. Thus, the JUNO-TAO experiment has been designed to have the following additional scientific goals: 1) provide a benchmark measurement for nuclear databases; 2) provide increased reliability in measured isotopic antineutrino yields; 3) provide an opportunity of improving nuclear physics knowledge of neutron-rich isotopes; 4) search for light sterile neutrinos with a mass-scale around 1 eV; 5) verify the technology for reactor monitoring and nuclear safeguard applications.

Design of the JUNO-TAO Experiment

The conceptual design of JUNO-TAO is shown on Figure 1. The Central Detector (CD) will collect reactor antineutrinos using 2.8 tons of Gadolinium-doped Liquid Scintillator (GdLS) inside a spherical acrylic vessel with a diameter of 1.8 m. To properly contain the energy deposition of gammas from the Inverse Beta Decay (IBD) positron annihilation, a position cut will be applied to remove events whose vertices are within 25 cm from the acrylic wall, resulting in a 1 ton fiducial mass. The IBD event rate in the fiducial volume is expected to be about 2000 (4000) events per day with (without) the detection efficiency taken into account. To reach the desired energy resolution, more than 4000 high-performance silicon photomultiplier (SiPM) tiles, each of dimensions 5 cm \times 5 cm, cover ~95% of the acrylic vessel and collect scintillation light with > 50% efficiency. The SiPM tiles are installed on the inner surface of a spherical copper shell with an inner diameter of 1.882 m. The copper shell is installed in a cylindrical stainless steel tank with a 2.1 m outer diameter and 2.2 m height. The stainless steel tank is filled with Linear Alkylbenzene (LAB), which serves as solvent for the GdLS and as as a buffer to shield the target from the outer tank's radioactivity. The LAB also allows to stabilize the temperature, and to optically couple the acrylic vessel to the SiPMs. Cooling pipes are deployed on the outer side of the copper shell and the inner side of the stainless steel tank. They are used to cool down the CD to -50° C, reducing the dark noise of SiPMs to $\sim 100 \text{ Hz/mm}^2$ and significantly suppressing its impact on the energy resolution. The stainless steel tank is insulated with 20 cm thick Polyurethane (PU) to reduce heat leakage. The central detector is surrounded by a 1.2 m thick water tank on the sides and 1 m High Density Polyethylene (HDPE) on the top as a shield against ambient radioactivity and cosmogenic neutrons. Cosmic muons are tagged by plastic scintillator strips on the top and the photomultiplier tubes that instrument the water tanks.

A photoelectron yield of about 4500 per MeV is expected from simulations, corresponding to an energy resolution of $1.5\%/\sqrt{E[{\rm MeV}]}$ in photoelectron statistics. When approaching the energy resolution limit of LS detectors, non-stochastic effects become prominent. At low energies, the contribution from the LS quenching effect could be quite large. At high energies, the smearing from neutron recoil of IBD becomes dominant. Taking into account the projected dark noise, cross-talk, and charge resolution of the SiPMs, the expected energy resolution of TAO will be sub-percent in most of the energy region of interest.

The muon and cosmogenic neutron rates in the JUNO-TAO site have been measured to be 1/3 of what they are in the surface. Simulations show that the accidental and cosmogenic backgrounds like fast neutrons and ${}^{8}\text{He}/{}^{9}\text{Li}$ decays can be controlled to < 10% of the signal with proper shielding and muon vetos.



Figure 1: Schematic view of the TAO detector, which consists of a liquid scintillator target with outer shielding and a veto system. Dimensions are displayed in mm.

Conclusions

In summary, the JUNO-TAO experiment is an important part of the JUNO experiment, and a leading player in the worldwide effort to precisely characterize the spectral shape of reactor antineutrinos. The detector R&D started in 2018. A recipe for GdLS has been developed and has been shown to have good transparency and light yield at -50° C. The SiPMs and the readout electronics have been preliminary tested at the same temperature. A first prototype with 100 L of GdLS has been used to verify the cryostat design. A fullsize prototype detector will be tested in 2021. A conceptual design report was released in 2020 [7]. The full JUNO-TAO detector is expected to start operations in 2022, around the same time as the larger JUNO experiment.

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